

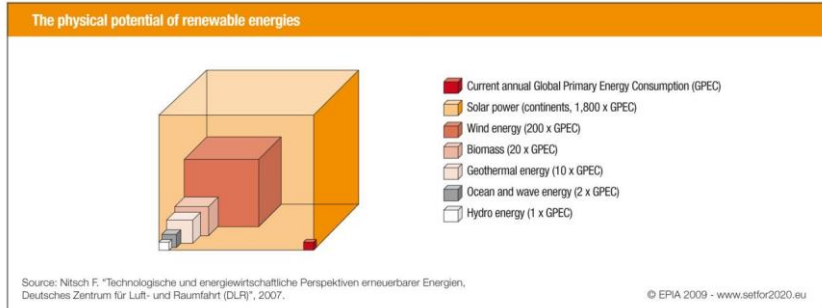


Solar resource

- Radiation from the Sun
- Atmospheric effects
- Insolation maps
- Tracking the Sun
- PV in urban environment

Solar resource

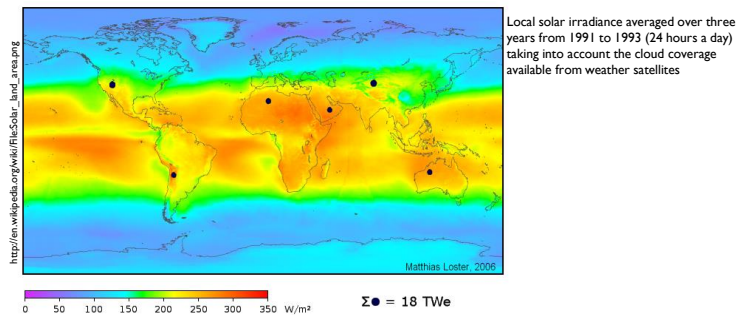
- Solar resource is immense
 - Human energy use: 4.0×10^{14} kWh/year
 - Solar resource on Earth's surface: 5.5×10^{17} kWh/year



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Solar resource

- Solar resource is immense
 - Human energy use: 4.0×10^{14} kWh/year
 - Solar resource on Earth's surface: 5.5×10^{17} kWh/year



Solar power systems covering the areas defined by the dark disks could provide more than the world's total primary energy demand (assuming a conversion efficiency of 8%).

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Solar resource

- Solar resource is **immense**
 - Human energy use: 4.0×10^{14} kWh/year
 - Solar resource on Earth's surface: 5.5×10^{17} kWh/year
- Solar resource is **global** and **democratic**
- Solar resource is relatively **constant** but depends on
 - atmospheric effects, including absorption and scattering
 - local variations in the atmosphere, such as water vapour, clouds, and pollution
 - latitude of the location
 - the season of the year and the time of day

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Solar resource

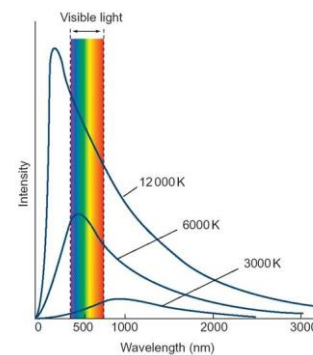


$$P_0 = \frac{\sigma T^4}{4\pi R_{sun}^2}$$

Total radiative power (Stefan Boltzman) $T=5762K$

Surface area of sun

Power radiated per unit area
 $5.96 \times 10^7 \text{ W/m}^2$



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Solar resource



$$P_0 = \frac{\sigma T^4}{4\pi R_{sun}^2}$$

Ratio of surface areas of the 2 spheres

Solar constant average energy flux incident at the Earth's orbit: **1366 W/m²**

$$S = \frac{4\pi R_{sun}^2}{4\pi D^2} P_0$$



Distance Sun-Earth

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Solar resource



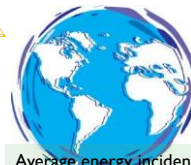
$$P_0 = \frac{\sigma T^4}{4\pi R_{sun}^2}$$

R_{sun} 6.96x10⁵ km
 D_{avg} 1.5x10⁸ km
 R_{Earth} 6.35x10³ km

$$S = \frac{4\pi R_{sun}^2}{4\pi D^2} P_0$$

Energy incident on Earth
 Total area of Earth

$$\frac{S \times \pi R_{Earth}^2}{4\pi R_{Earth}^2} = \frac{S}{4}$$

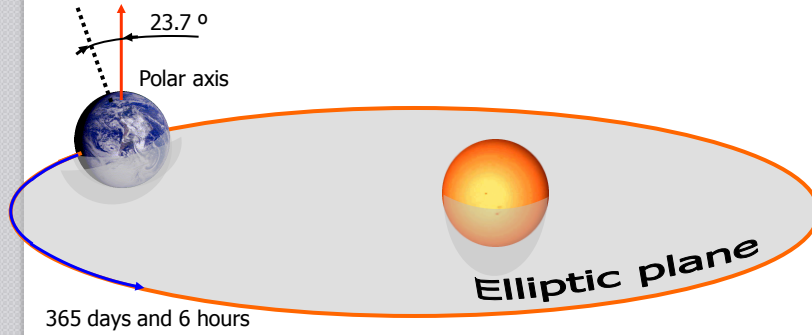


Average energy incident per unit area of surface of Earth: **342 W/m²**

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Solar resource

- Earth-Sun motion



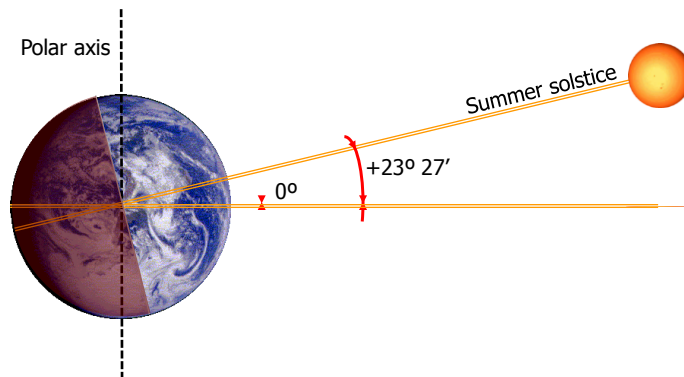
$$\frac{H}{S} = 1 + 0.033 \cos\left(\frac{360(n-2)}{365}\right)$$

$H(W/m^2)$ is radiant power density outside the atmosphere; S is solar constant; n is day of the year

Solar resource

- Earth-Sun motion

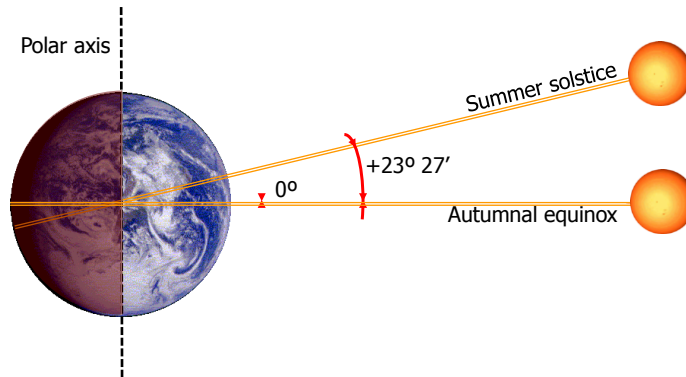
- **Solar declination:** angle between line joining centres of Earth and Sun and the equatorial plane



Solar resource

- Earth-Sun motion

- **Solar declination:** angle between line joining centres of Earth and Sun and the equatorial plane

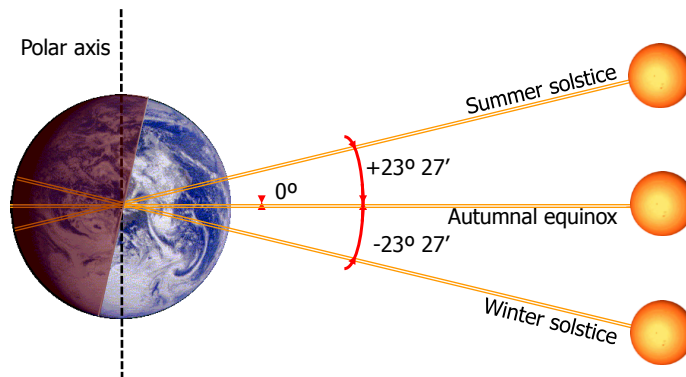


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Solar resource

- Earth-Sun motion

- **Solar declination:** angle between line joining centres of Earth and Sun and the equatorial plane

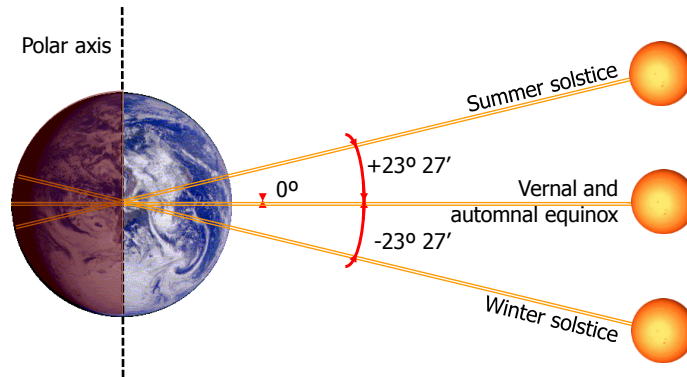


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Solar resource

- Earth-Sun motion

- **Solar declination:** angle between line joining centres of Earth and Sun and the equatorial plane

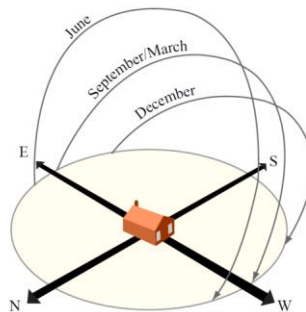


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Solar resource

- Earth-Sun motion

- **Solar declination:** angle between line joining centres of Earth and Sun and the equatorial plane



Building orientation with the long axis facing south

$$\delta = \pi \frac{23.45}{180} \sin\left(2\pi \frac{284+n}{365}\right)$$

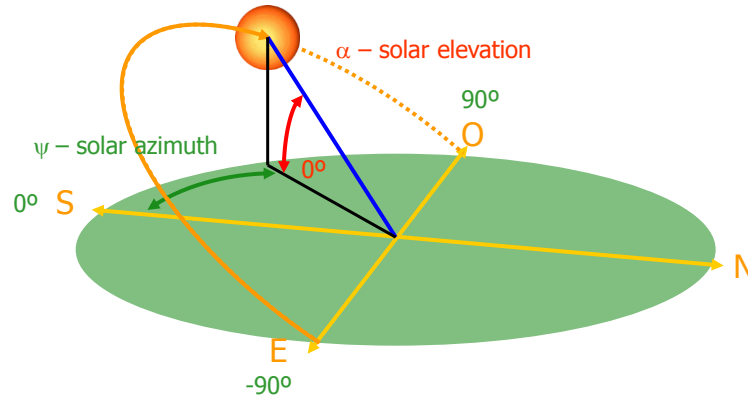
Declination in radians; n is the number of the day (Jan 1st = 1)

Solar resource

- Earth-Sun motion

$$\sin \alpha = \sin \delta \sin \phi + \cos \delta \cos \phi$$

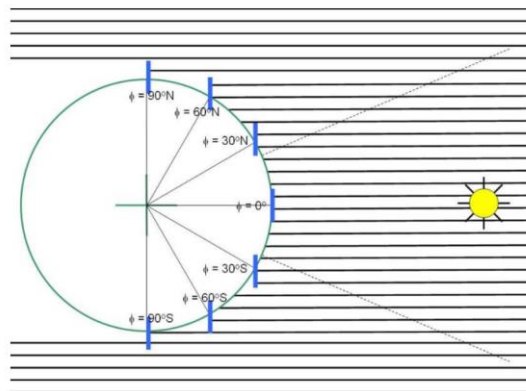
$$\cos \psi = \frac{\sin \alpha \sin \phi - \sin \delta}{\cos \alpha \cos \phi}$$



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Solar resource

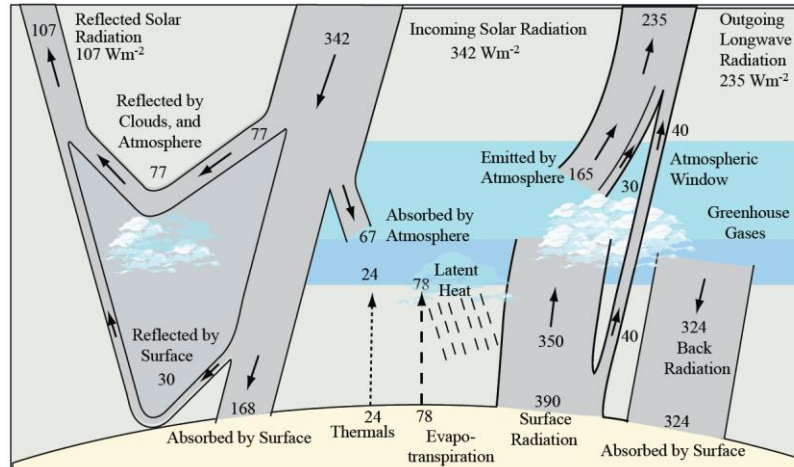
- Optimum orientation: facing south (north in the southern hemisphere)
- Optimum inclination: local latitude – but not quite



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Solar resource

Atmospheric effects



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Solar resource

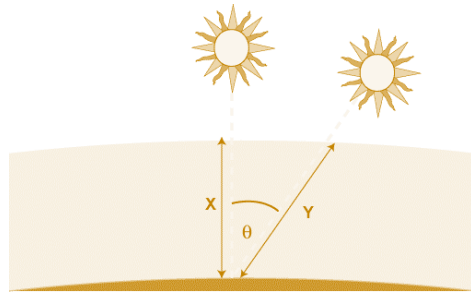
Atmospheric effects on solar radiation at the Earth's surface:

- a **reduction in the power** of the solar radiation due to absorption, scattering and reflection in the atmosphere;
- a **change in the spectral content** of the solar radiation due to greater absorption or scattering of some wavelengths;
- the **introduction of a diffuse** or indirect component into the solar radiation; and
- local variations in the atmosphere (such as water vapour, clouds and pollution) which have additional effects on the incident power, spectrum and directionality.

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Solar resource

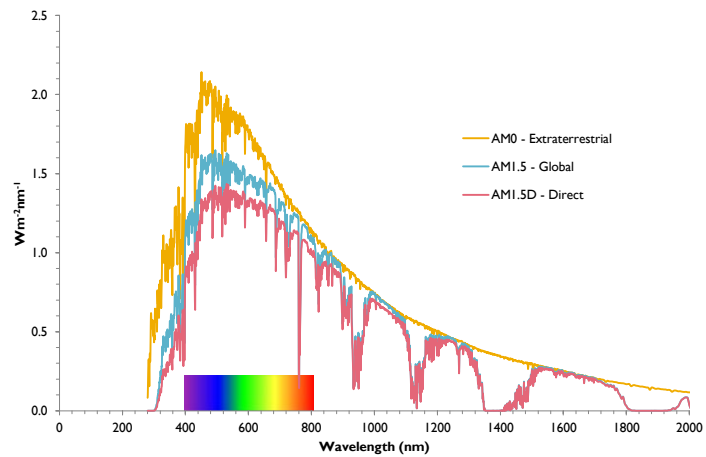
- **Air Mass** is a measure of the reduction in the power of light as it passes through the atmosphere and is absorbed by air and dust



$$AM = \frac{1}{\cos \theta}$$

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Solar resource



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Solar resource

- Photon flux Φ [photons/s/m²]
- Power density H [W/m²]
- Spectral irradiance F [W/m²/μm] is the power density at a given λ (μm)

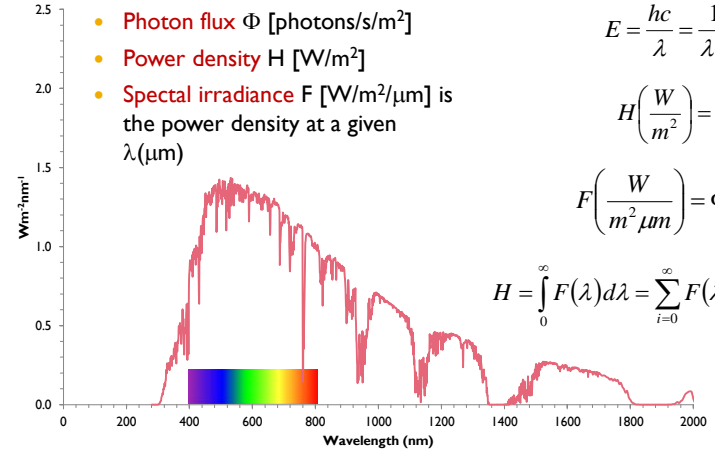
$$\Phi = \frac{\text{photons}}{m^2 s}$$

$$E = \frac{hc}{\lambda} = \frac{1.24}{\lambda(\mu m)}$$

$$H\left(\frac{W}{m^2}\right) = \Phi \frac{hc}{\lambda}$$

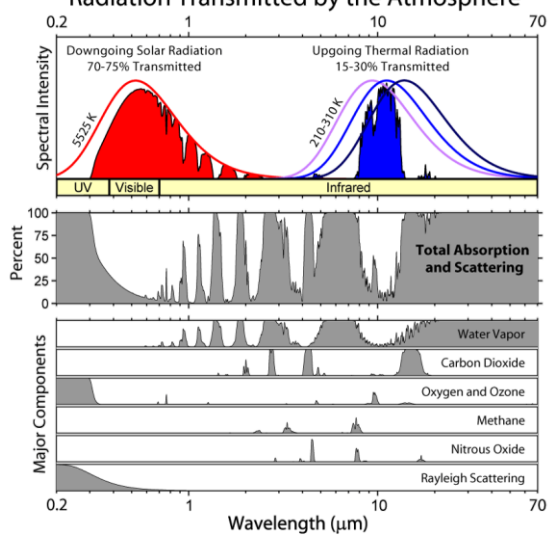
$$F\left(\frac{W}{m^2 \mu m}\right) = \Phi \frac{hc}{\lambda^2}$$

$$H = \int_0^{\infty} F(\lambda) d\lambda = \sum_{i=0}^{\infty} F(\lambda) \Delta\lambda$$

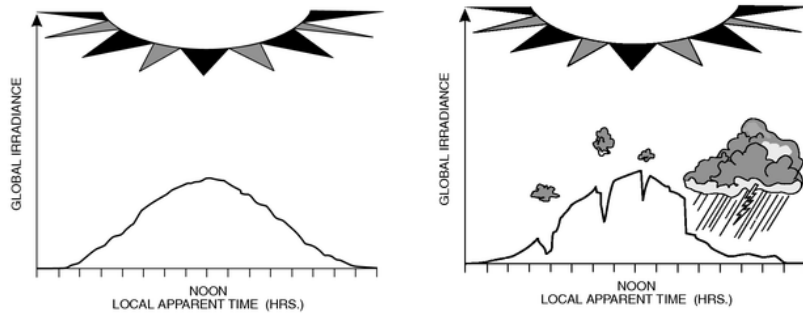


Solar resource

Radiation Transmitted by the Atmosphere



Solar resource



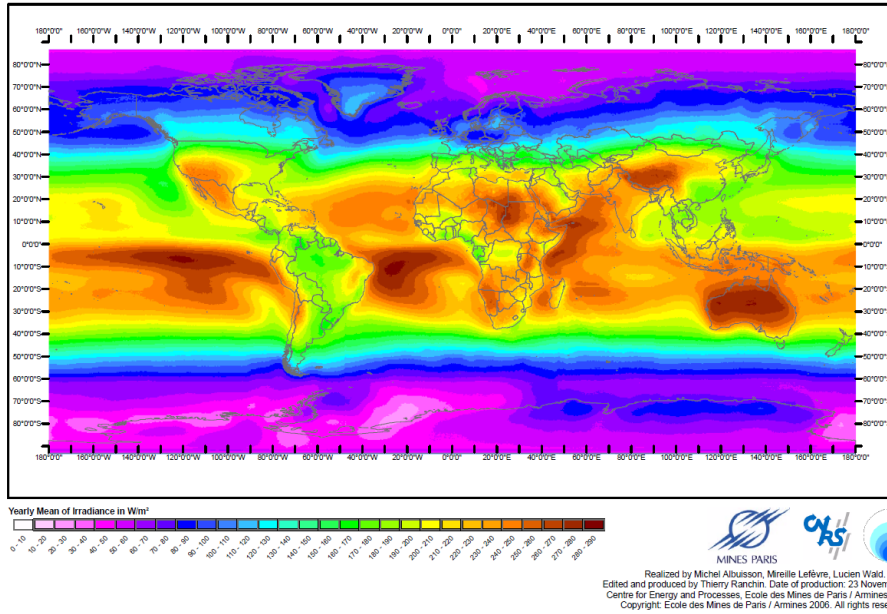
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Solar resource

- Insolation: **Incoming Solar Radiation**
- Typical units: **kWh/m²/day**
- Affected by latitude, local weather patterns,...
- Šuri M., Huld T.A., Dunlop E.D. Ossenbrink H.A., 2007. *Potential of solar electricity generation in the European Union member states and candidate countries*. Solar Energy, 81, 1295–1305, <http://re.jrc.ec.europa.eu/pvgis/>.

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Averaged Solar Radiation 1990-2004

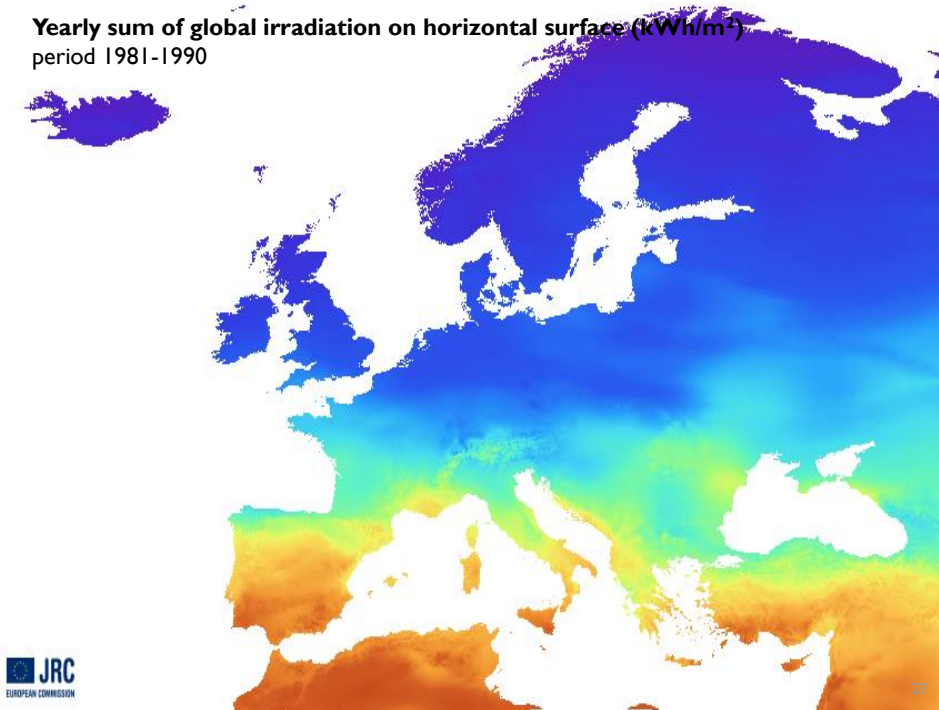


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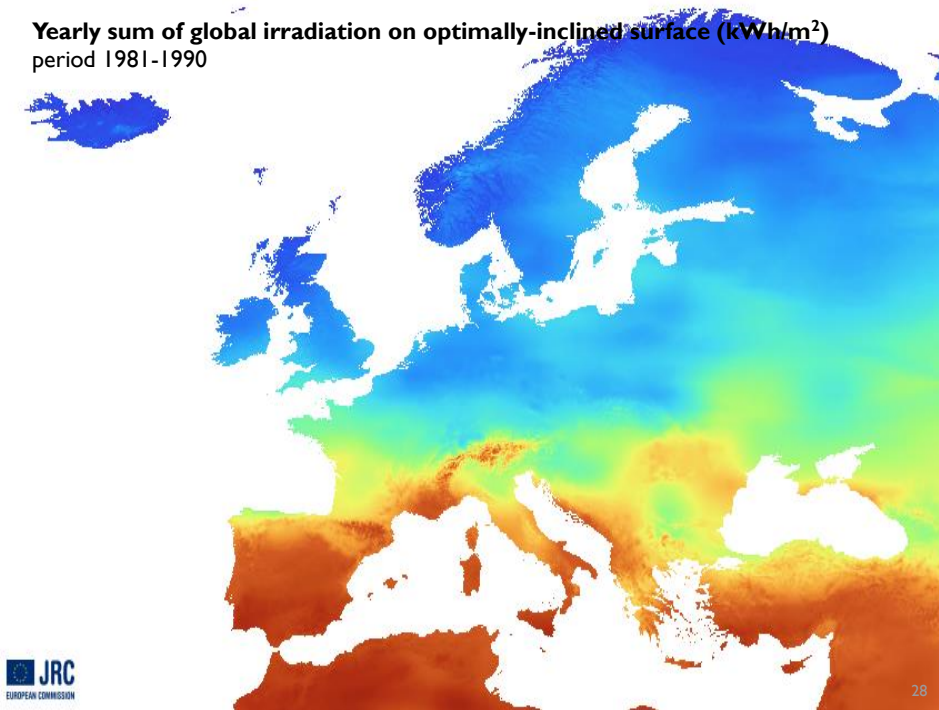
Yearly sum of global irradiation on vertical surface (kWh/m^2) period 1981-1990



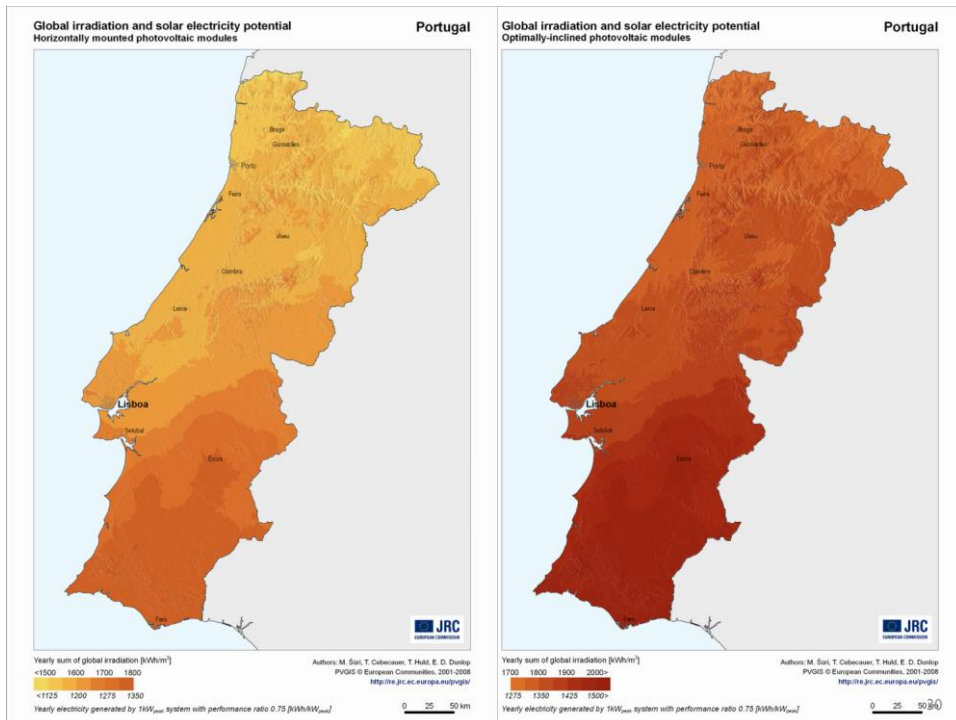
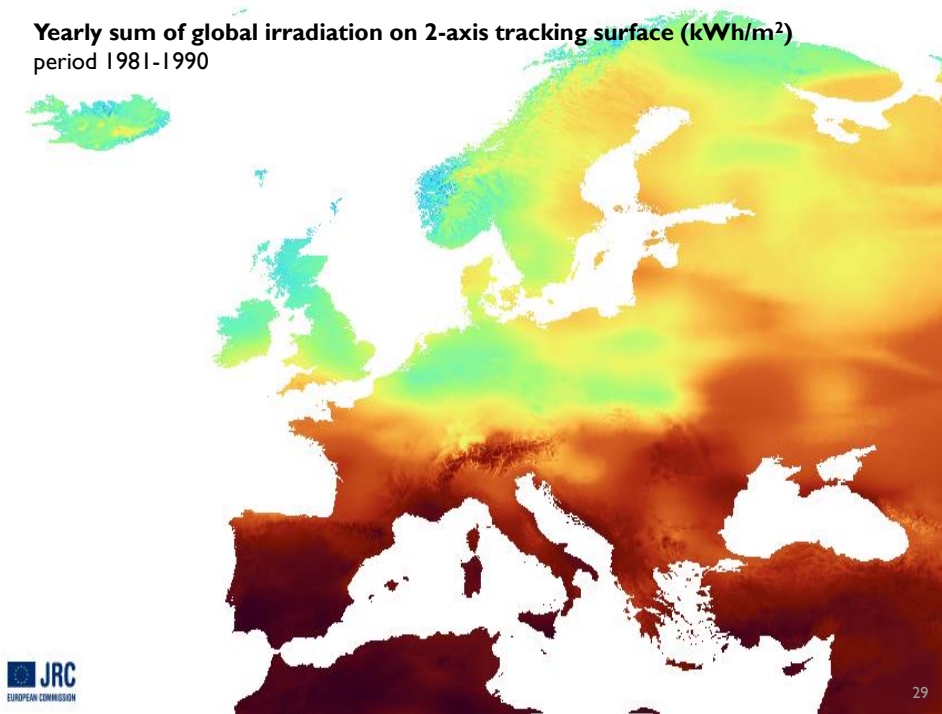
Yearly sum of global irradiation on horizontal surface (kWh/m²)
period 1981-1990



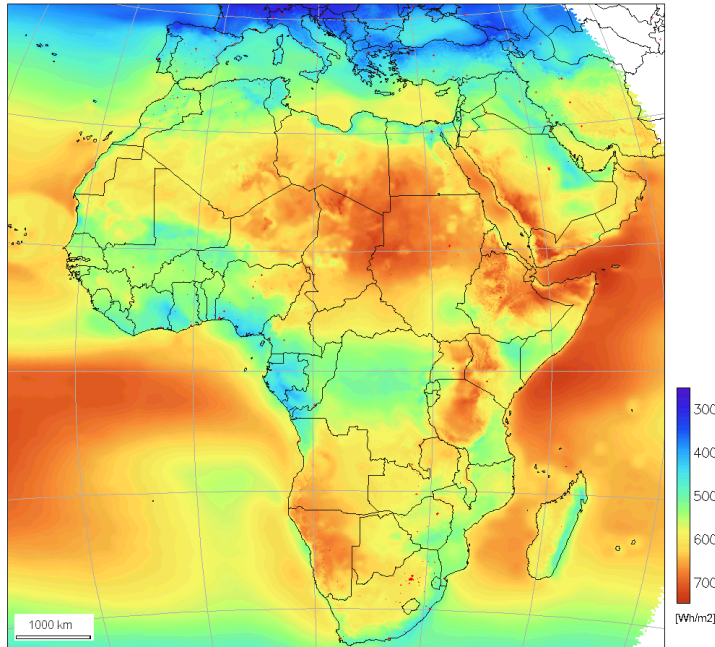
Yearly sum of global irradiation on optimally-inclined surface (kWh/m²)
period 1981-1990



**Yearly sum of global irradiation on 2-axis tracking surface (kWh/m²)
period 1981-1990**



Global horizontal irradiation (1985-2004)
(annual average of daily sums, Gh)



PVGIS (c) European Communities 2002-2008
HelioClim-1 (c) Ecole des Mines de Paris/ARMINES 1985-2005

<http://re.jrc.ec.europa.eu/pvgis/>

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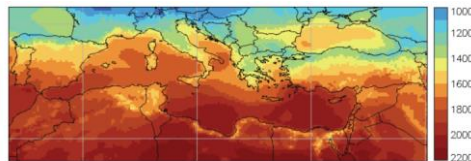


Figure 1: Long-term average of yearly sums of global horizontal irradiation (kWh/m², time series representing years 1985, 1987, and 1989-2004)

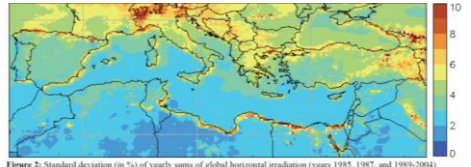


Figure 2: Standard deviation (in %) of yearly sums of global horizontal irradiation (years 1985, 1987, and 1989-2004)

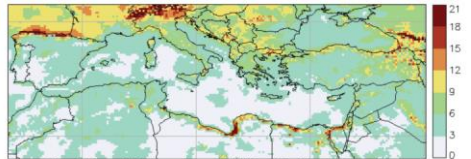


Figure 6: Relative difference of the highest yearly sum of global horizontal irradiation in relation to the long-term average (in %)

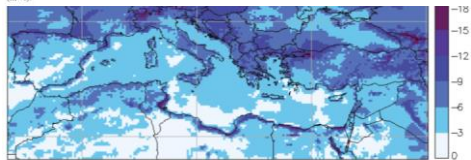


Figure 5: Relative difference of the lowest yearly sum of global horizontal irradiation in relation to the long-term average (in %)

Solar resource

- Coastal areas and higher mountains face wider variations (up to 10%)
- Winter is much more variable (up to x6) than summer months

Šúri M., Huld T., Dunlop E.D., Albuissou M., Lefèvre M., Wald L., 2007. *Uncertainties in photovoltaic electricity yield prediction from fluctuation of solar radiation*. Proceedings of the 22nd European Photovoltaic Solar Energy Conference, Milano, Italy 3-7.9.2007

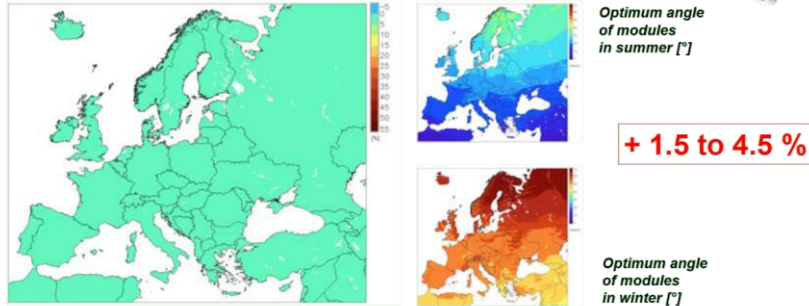
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Solar tracking

Compared to PV with modules fixed at optimum angle:

- Changing inclination twice a year contributes only marginally

Fixed mounting - two (seasonal) optimum angles

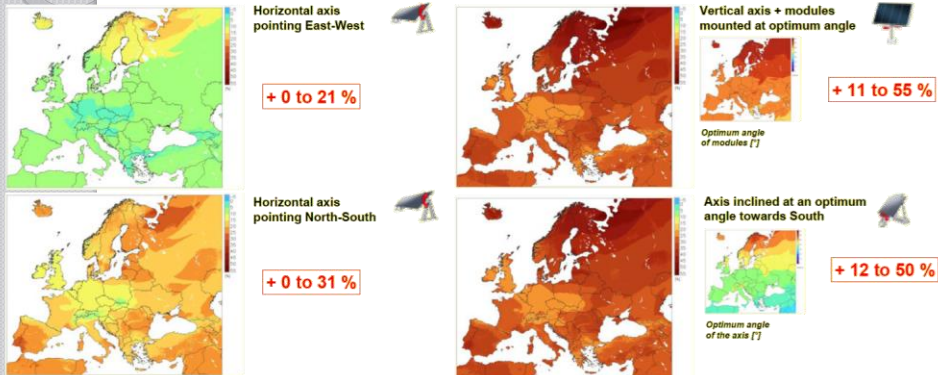


Comparison of electricity yield from fixed and suntracking PV systems in Europe [JRC, 2008]

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Solar tracking

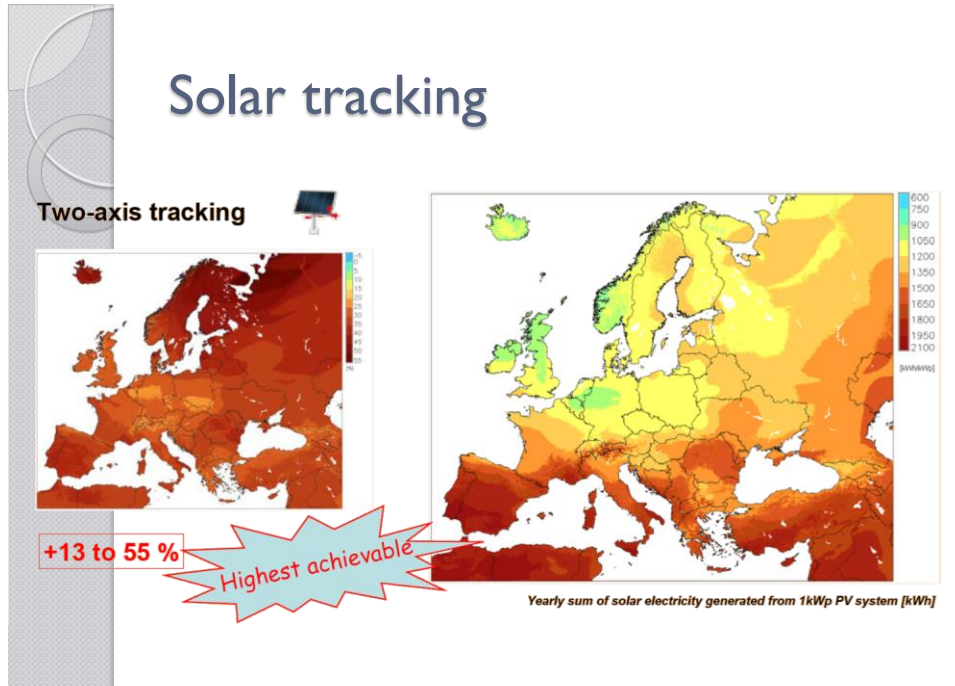
One-axis tracking



Comparison of electricity yield from fixed and suntracking PV systems in Europe [JRC, 2008]

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Solar tracking



Comparison of electricity yield from fixed and suntracking PV systems in Europe [JRC, 2008]

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Solar tracking

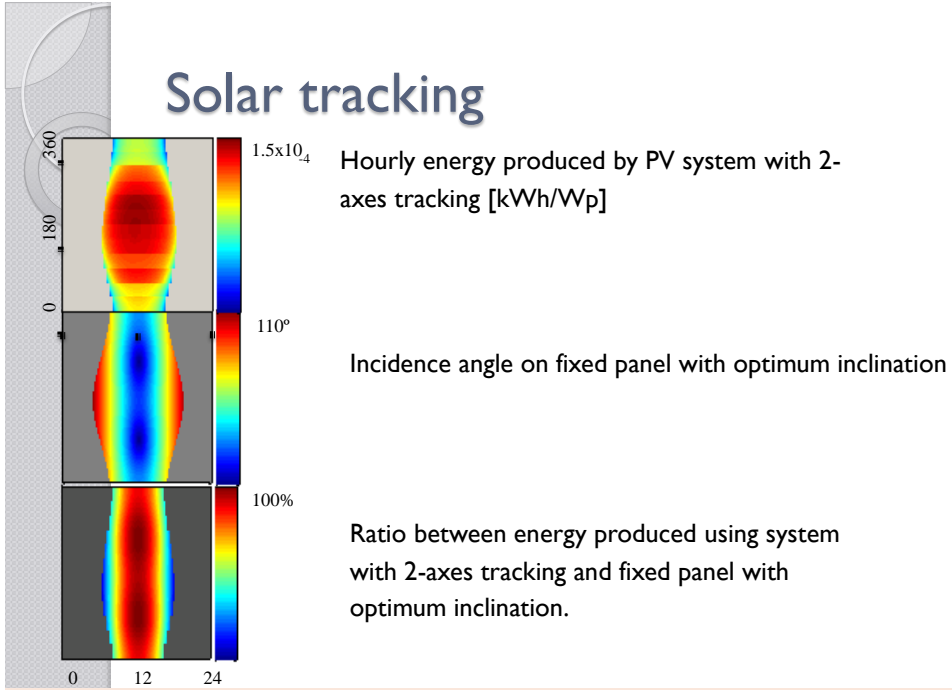
Compared to PV with modules fixed at optimum angle:

- Changing inclination twice a year contributes only marginally (2-4%)
- I-axis tracking PV with vertical or South-inclined axis generates only 1-4% less than 2-axis tracking system
- I-axis tracking PV with horizontal axis oriented E-W typically performs only slightly better than fixed mounting systems

Comparison of electricity yield from fixed and suntracking PV systems in Europe [JRC, 2008]

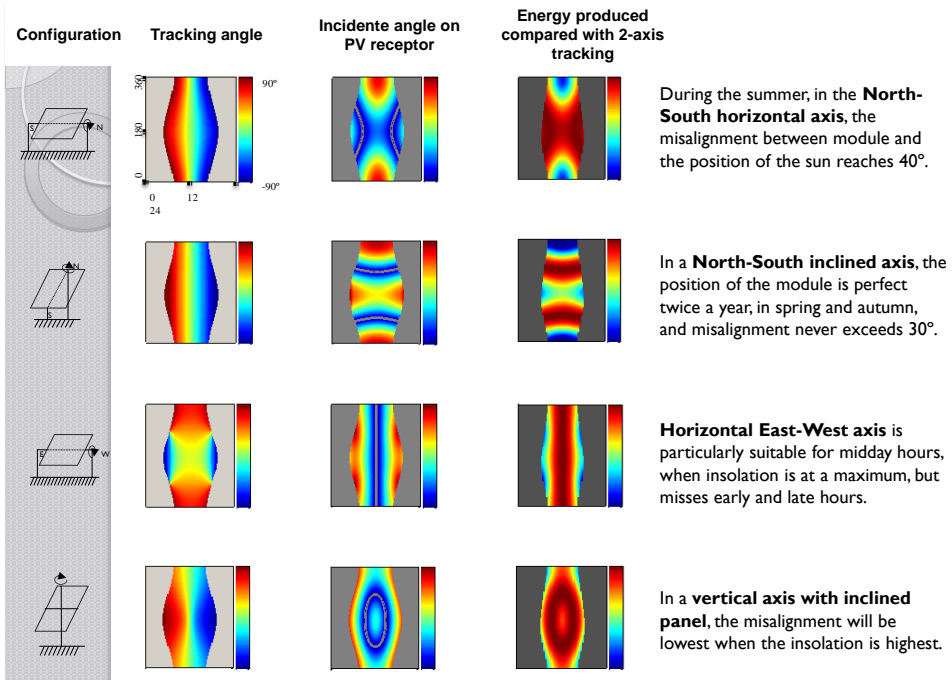
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Solar tracking



Gaspar et al, Exploring One-Axis Tracking Configurations For CPV Application, CPV7, Las Vegas 2011

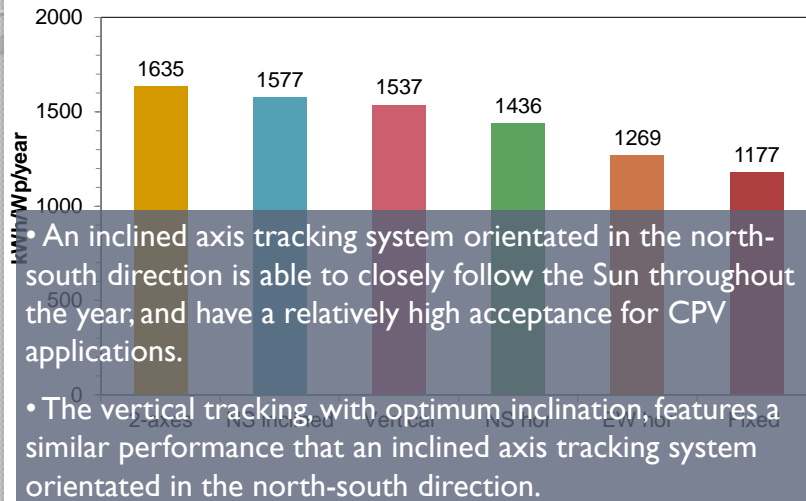
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Gaspar et al, Exploring One-Axis Tracking Configurations For CPV Application, CPV7, Las Vegas 2011

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Solar tracking



Gaspar et al, Exploring One-Axis Tracking Configurations For CPV Application, CPV7, Las Vegas 2011

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Solar tracking



▲ Installing a screw foundation for a dual-axis tracking system (Dager-Traker) at Am Peterswald PV park in Germany. Bftec GmbH pre-drills the ground at the site and then the foundation is installed. The foundation's unusual form, which consists of a smooth side wall with a coil at the bottom (if necessary, two coils), guarantees a firm grip. The foundation can rotate in the ground without displacing or loosening the surrounding earth.

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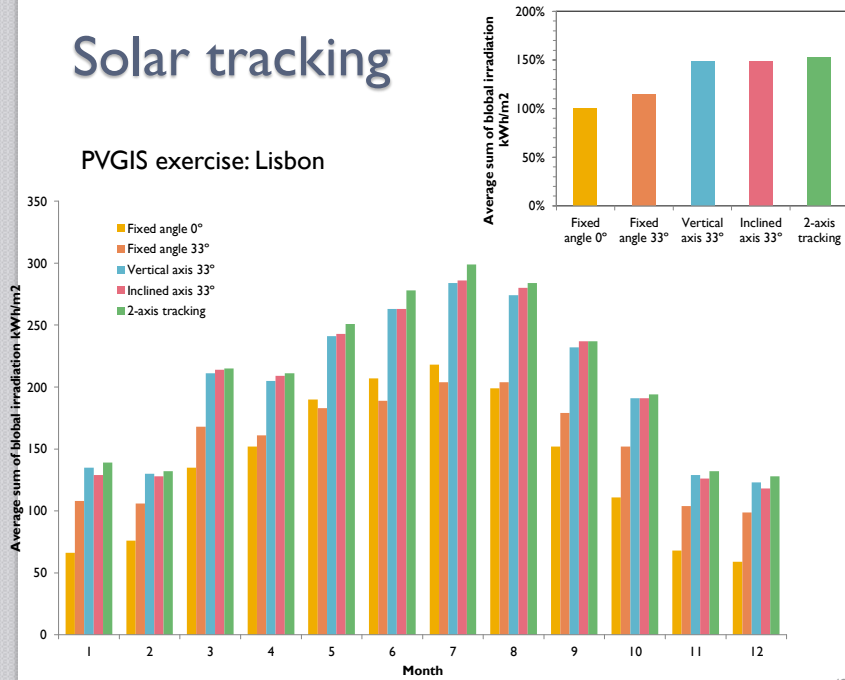
Solar tracking



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Solar tracking

PVGIS exercise: Lisbon



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Solar tracking

Shading effect

- Ground cover ratio = PV area / total area

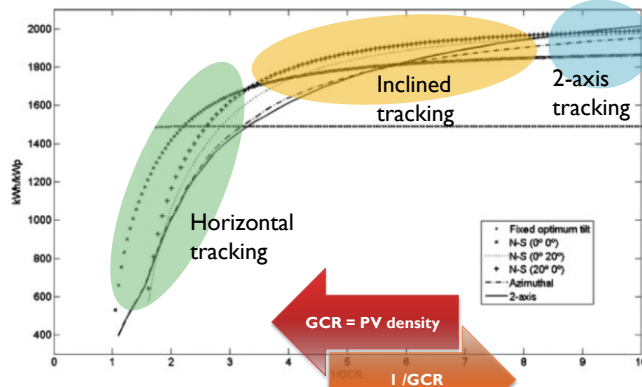


Figure 13. Evolution of yearly energy yield in kWh/kWp for the pessimistic shading case and assuming a constant dirtiness degree of 3%

E. Narvarde et al, Tracking and Ground Cover Ratio, Prog. Photovolt: Res. Appl. 2008; 16:703–714

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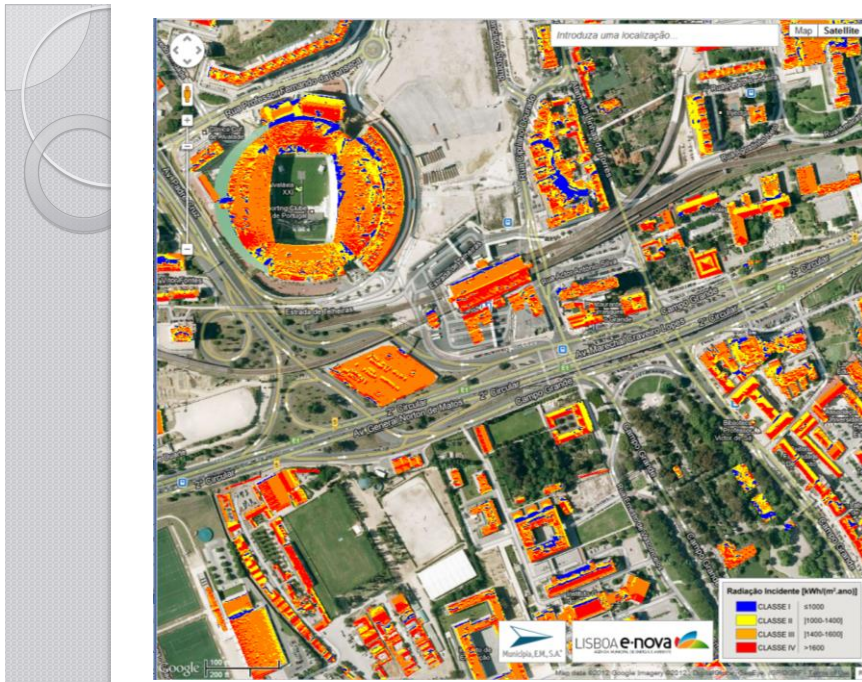
Solar in the city

PV potential in the urban landscape is harder to estimate

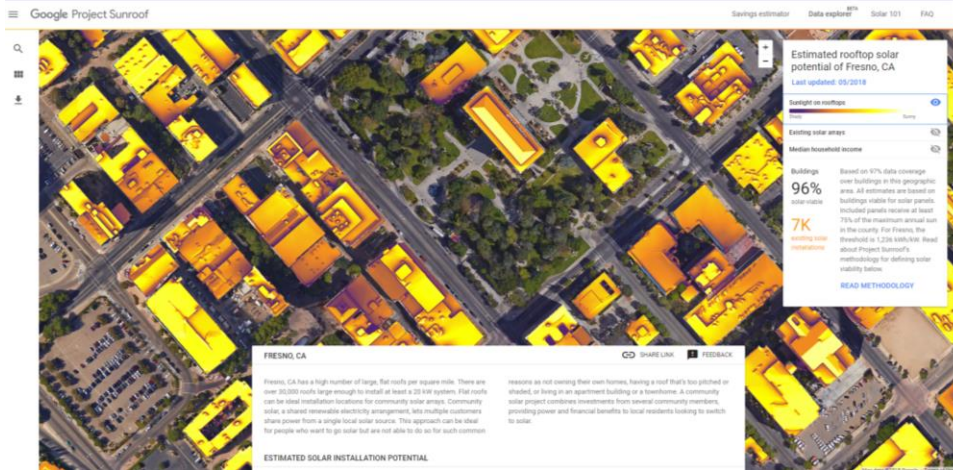
- **Geographical** solar potential
locations where this energy can be captured
- **Technical** solar potential
technical characteristics of the rooftop/equipment used
- **Economic** solar potential
only viable systems

<https://doi.org/10.1016/j.solener.2008.03.007>

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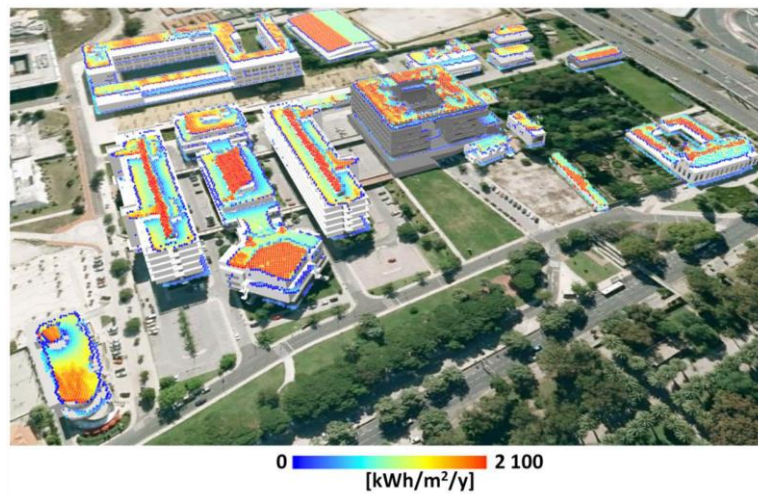


Solar in the city



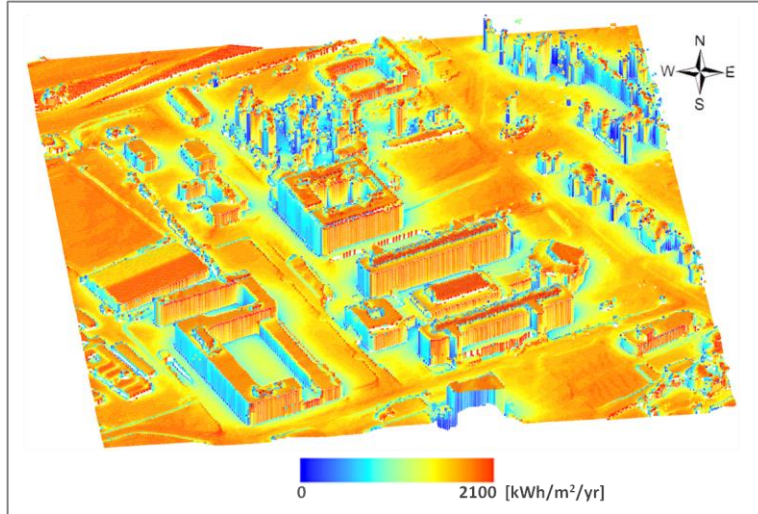
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Solar in the city



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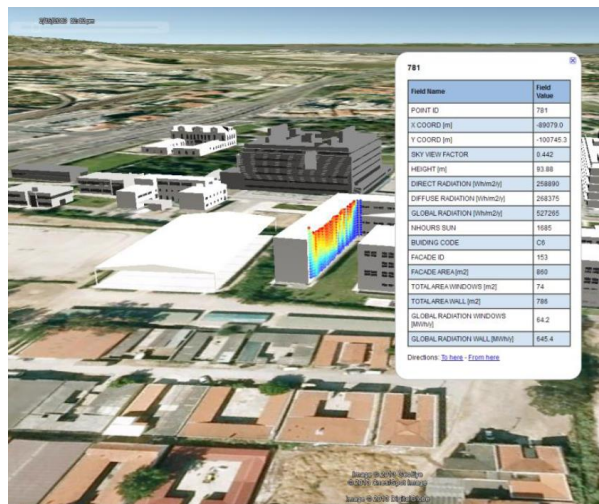
Solar in the city



Redweik et al. Solar Energy 97 (2013) 332–341

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Solar in the city



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